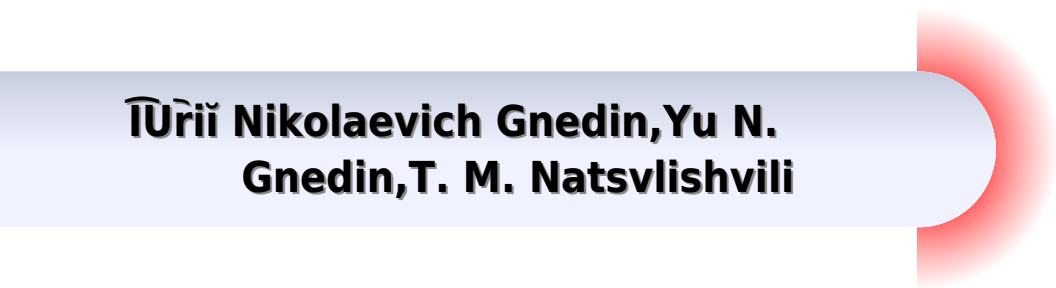


Object	P_c (yr)	A	M_2/M_1
HD 16160	12.3	1.0	0.02
HD 146233	8	1.0	0.14
HD 26913 (y)	5.8	1.0	0.17
HD 160346	7.1	1.2	0.03
HD 161239	5.1	1.3	0.02
HD 10476	10.4	1.3	0.16
HD 4628	8.5	1.4	0.02
Sun	11	1.5	0.03
HD 81809	8.1	1.5	0.08
HD 26965 (y)	10	1.5	0.18
HD 103095	7.2	1.5	0.18
HD 187691	8	1.7	0.08
HD 219834B	16	1.7	0.15
HD 20630 (y)	5.5	2.9	0.11
HD 18256(y)	7	3.8	0.27

Anharmonic And Standing Dynamo Waves Theory And Observation Of Stellar Magnetic Activity

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Gnedin, T. M. Natsvlshvili**



Anharmonic And Standing Dynamo Waves Theory And Observation Of Stellar Magnetic Activity:

Anharmonic and Standing Dynamo Waves: Theory and Observation of Stellar Magnetic Activity, 2005 The familiar decadal cycle of solar activity is one expression of interannual variability of surface magnetism observed in stars on or near the lower main sequence Records of Ca II H and K fluxes obtained for such stars by Mount Wilson Observatory's HK Project extend back over 35 years From these records we define a new measure of anharmonicity that connects to the 2 D description of a Parker dynamo model We explore the parameter space of this model and find an excellent counterpart to solutions containing highly anharmonic standing dynamo waves in the records of several of the lowest mass late K to early M type active stars in the sample We interpret anharmonicity as resulting from non propagating or standing dynamo waves which operate in a substantially supercritical regime There for most of the cycle large scale magnetic fields are generated and maintained by winding of field by differential rotation rather than by joint action of differential rotation and helical convection Further less active stars like the Sun show simple harmonic migratory and or intermediate type dynamo wave patterns over a broad range of dynamo parameters Energy Research Abstracts, 1987 An Introduction to Stellar Magnetic Activity BASRI, Professor Gibor Basri, 2021-10-31 The book provides an introduction to the observational and theoretical knowledge of solar and stellar magnetic activity for graduate or advanced undergraduate students **Stellar Magnetic Activity** Carolus J. Schrijver, 1986 **Magnetic Fields of Stars** Ю́рий Николаевич Гнедин, Yu N. Gnedin, T. M. Natsvlishvili, 2000 **Stellar Magnetism** Leon Mestel, 2012-02-16 Ongoing studies in mathematical depth and inferences from helioseismological observations of the internal solar rotation have shown up the limitations in our knowledge of the solar interior and of our understanding of the solar dynamo manifested in particular by the sunspot cycle the Maunder minimum and solar flares This second edition retains the identical overall structure as the first edition but is designed so as to be self contained with the early chapters presenting the basic physics and mathematics underlying cosmical magnetohydrodynamics followed by studies of the specific applications appropriate for a book devoted to a central area in astrophysics *Dynamo and Dynamics, a Mathematical Challenge* Pascal Chossat, Dieter Armbruster, Iuliana Oprea, 2012-12-06 This book contains the lectures given at the workshop Dynamo and dynamics a mathematical challenge held in Cargèse from August 21 to 26 2000 The workshop differed from most previous conferences on the dynamo effect in two important respects First it was at this international conference that the experimental observation of homogeneous fluid dynamos was first reported Second the conference gathered scientists from very different fields thus showing that the dynamo problem has become an interdisciplinary subject involving not only astrophysicists and geophysicists but also scientists working in dynamical systems theory hydrodynamics and numerical simulation as well as several groups in experimental physics This book thus reports important results on various dynamo studies in these different contexts Decades after the discovery of the first analytic examples of laminar fluid dynamos the self generation of a magnetic field by a flow

of liquid sodium has been reported by the Karlsruhe and Riga groups. Although there were no doubts concerning the self-generation by the laminar Roberts type or Ponomarenko type flows that were used, these experiments have raised interesting questions about the influence of the turbulent fluctuations on the dynamo threshold and on the saturation level of the magnetic field.

Magnetic Fields and Star Formation, 2004 *Solar and Astrophysical Dynamos and Magnetic Activity (IAU S294)* Alexander G. Kosovichev, Yihua Yan, Elisabete de Gouveia Dal Pino, Lidia van Driel-Gesztelyi, 2013-09-12

Global magnetic fields in planets in the Sun and other stars in spiral galaxies and galaxy clusters are believed to be generated and maintained by a hydromagnetic dynamo, a process that converts turbulent kinetic energy into magnetic energy. These dynamo processes operate on drastically different scales but are associated with common physical mechanisms involving a complex interaction of rotation, turbulence, and instabilities. The goal of IAU Symposium 294 was to discuss the most important results of recent studies of the cosmic dynamo processes from planets to stars, galaxies, and clusters of galaxies. This volume covers advances in dynamo theories and numerical simulations, links between dynamos and turbulence, the origin of magnetic fields, and current and future observational projects. The proceedings of IAU S294 are an important asset for advanced students and researchers as a summary of hot topics related to the solar and astrophysical dynamos and magnetic activity.

Stellar and Solar Magnetic Fields R. Lüster, 1965

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